

NUCLEAR CHEMISTRY

Nuclear reactions

reactions involve atomic nuclei

lead to changes in atoms

conversion of an atom into another

Isotopes

atoms having same atomic number

but different mass number

oxygen-17

Nuclear reactions

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conversion of an atom into another

Isotopes

atoms having same atomic number

but different mass number



Two sources of radioactivity

Natural radioactivity

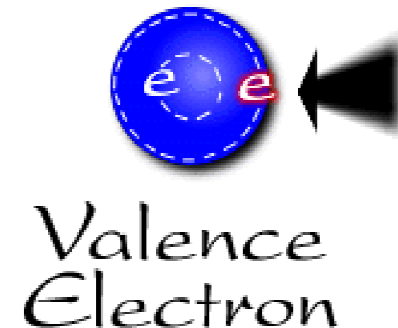
isotopes present since the earth formed **eg uranium-238 or 235**
or produced by cosmic rays from the sun **eg carbon-14**

Synthetic radioactivity

isotopes made in nuclear reactors when atoms are split (fission).
Cyclotrons, particle accelerators

What is radioactivity?

- ▶ Normal chemical reactions involve electrons **Valence electrons**



What is radioactivity?

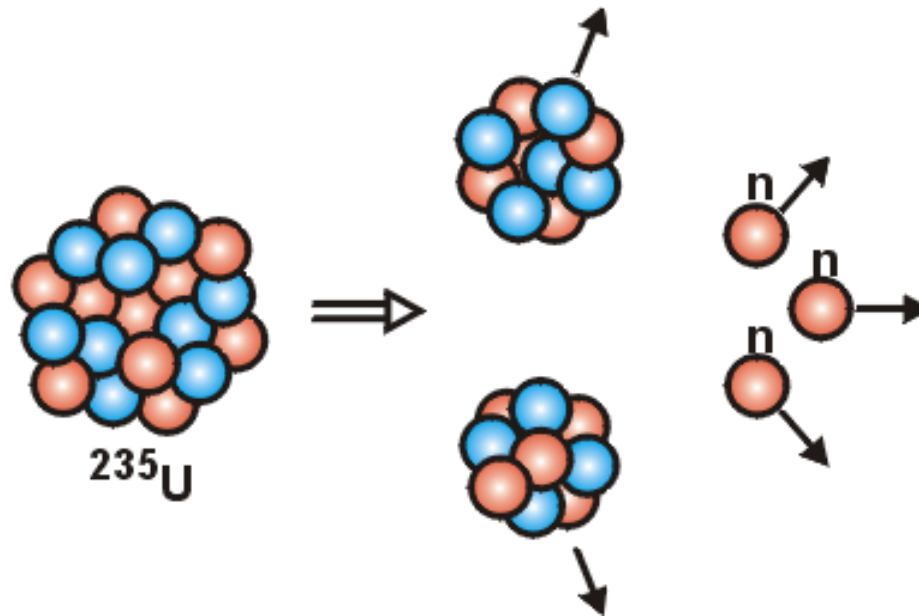
- ▶ **unstable atoms** \Rightarrow **stable atoms**

H																	He
Li	Be											B	C	N	O	F	Ne
Na	Mg											Al	Si	P	S	Cl	Ar
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
Cs	Ba	Lu	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
Fr	Ra	Ac	Rf	Db	Sg	Bh	Hs	Mt	110	111	112						
La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy	Ho	Er	Tm	Yb				
Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	Md	No				

Elements that are naturally radioactive

What is radioactivity?

- ▶ **unstable atoms** \Rightarrow **stable atoms**



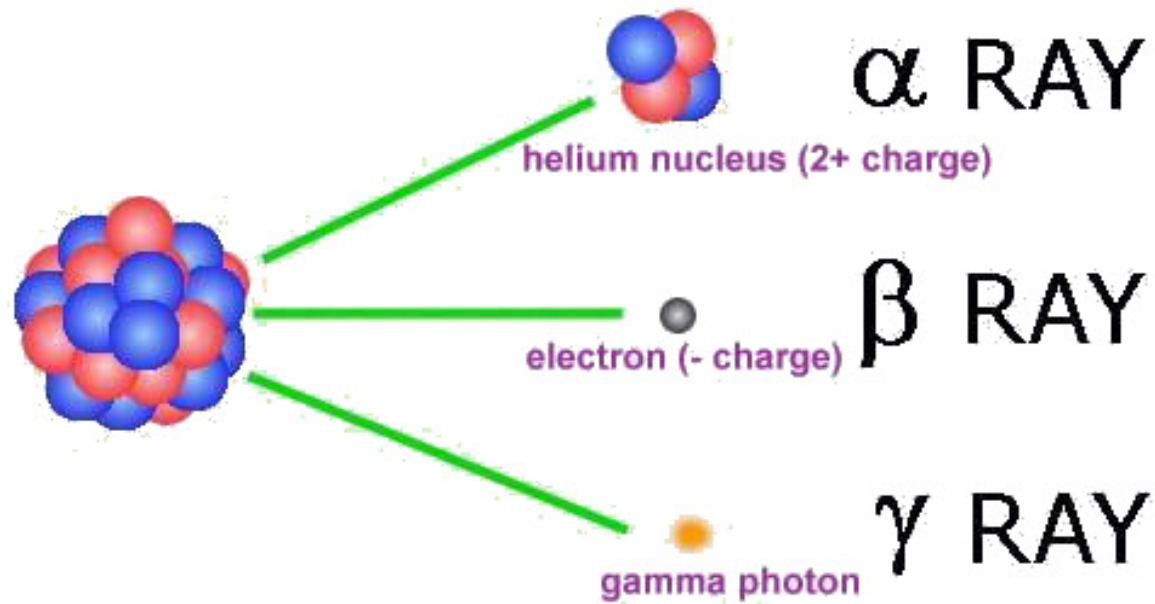
What is radioactivity?

Our understanding of atomic structure came from studies of radioactive elements

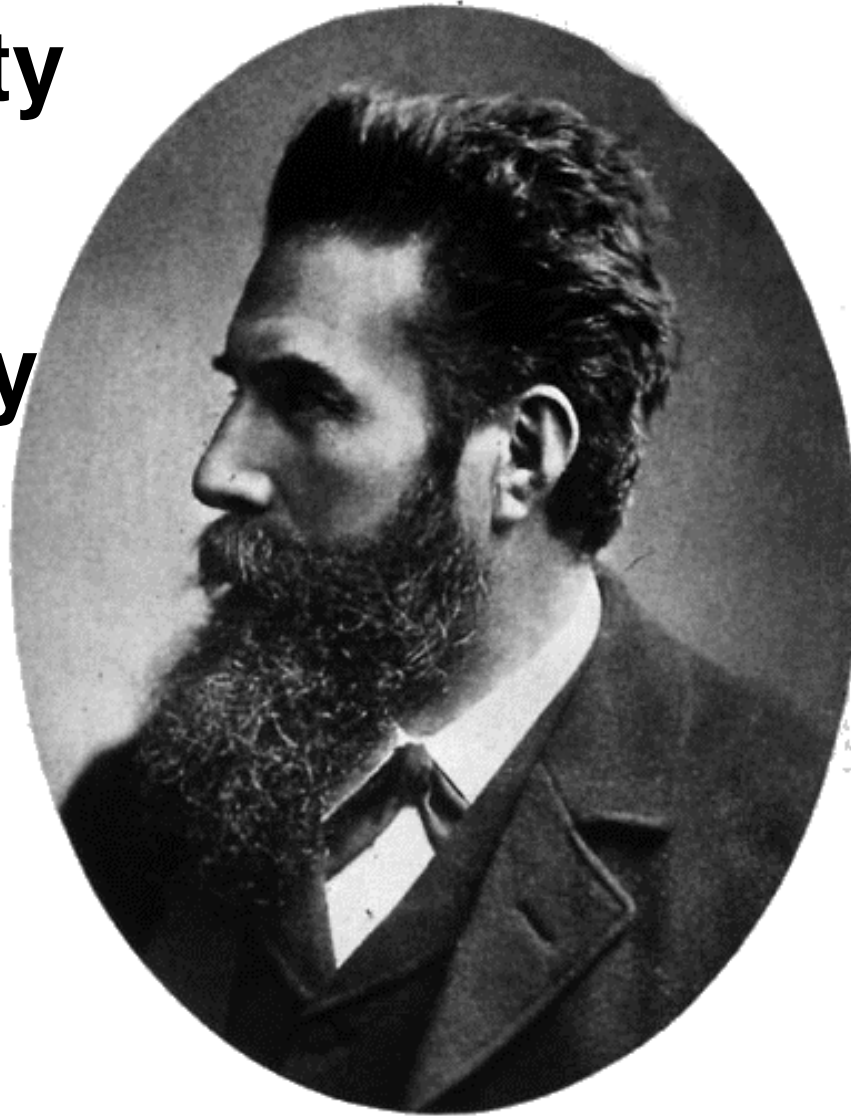
What is radioactivity?

the process by which atoms spontaneously emit high energy particles or rays from their nucleus

What is radioactivity?



**Radioactivity
unknown
before
20th century**



**Discovered
by
Röntgen
and
Becquerel
Röntgen
discovered
X-rays in
cathode
ray tubes
in 1895**

Wilhelm Röntgen (1845-1923)

**Röntgen: rays produced a fog on
photographic plates**

**World's first
X-ray:
Bertha
Röntgen's
hand**



Becquerel (1845-1923)

U compounds

produced fog on

photographic plates

Discovered (1896) it was

naturally radioactive



1867-1934

**Showed other
elements -
thorium- were
radioactive**

**With Pierre Curie,
discovered and
isolated new
elements:
radium, polonium**

**Radium millions of
times more active
than uranium**

Radium

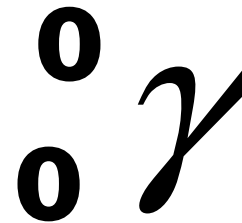
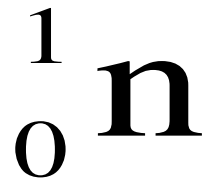
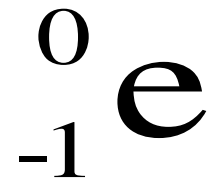


**Good for nervous disorders,
insomnia, general debility,
arthritis, and rheumatism.**

**This large pottery crock
was lined with radium ore**

**Vita Radium Suppositories
(ca.1930)**

Common particles



Types of Nuclear Reactions

1. Radioactive decay:
unstable nucleus

2. Transmutation:
Atoms change into new atoms

Types of Nuclear Reactions

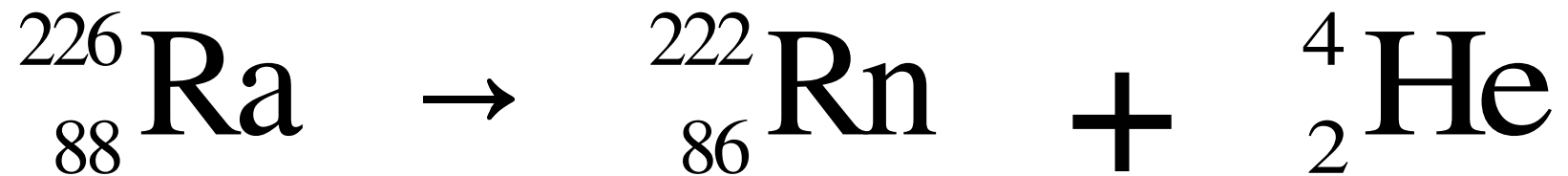
3. Fission:

Heavy into lighter nucleus

4. Fusion

Light into heavier nucleus

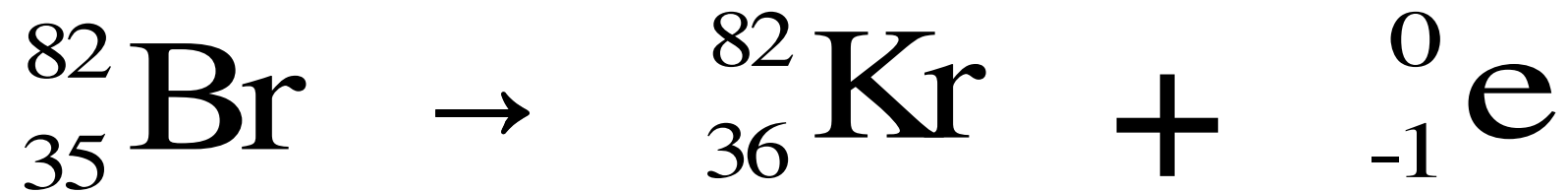
Nuclear Equations



Atomic numbers must equal

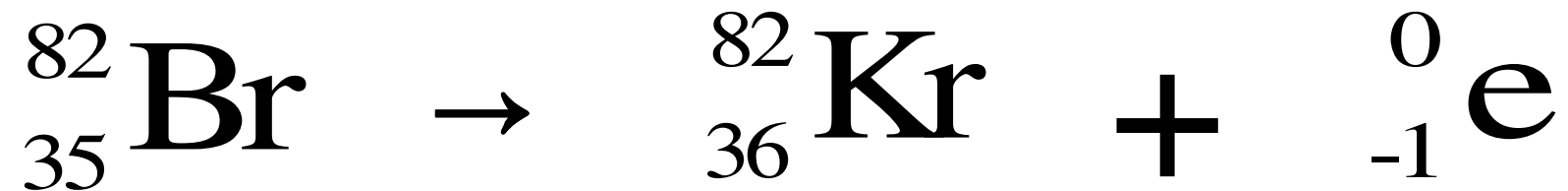
Mass numbers must equal

Nuclear Equations

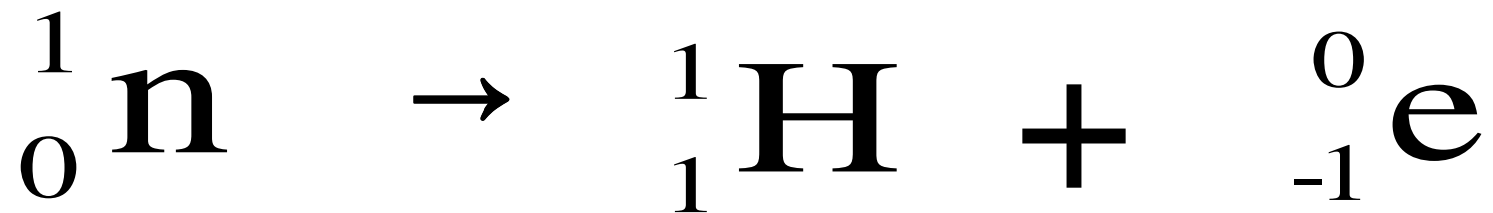


How can electrons come from nucleus?

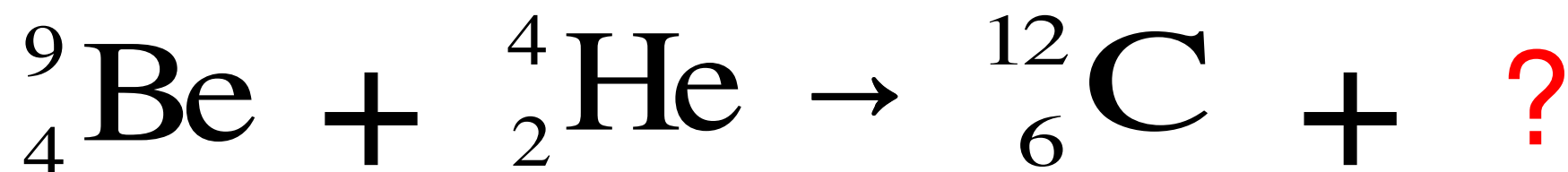
Nuclear Equations



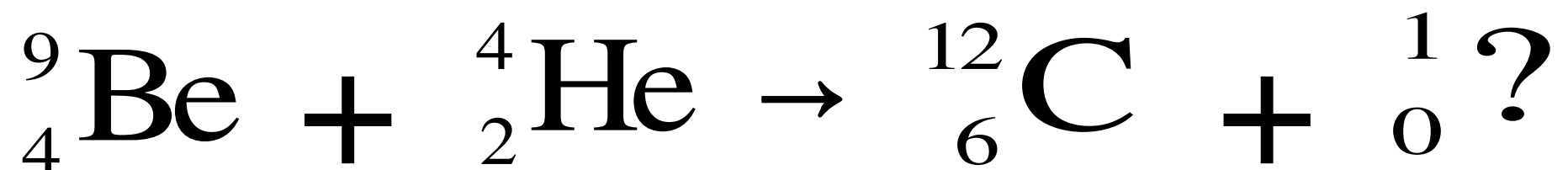
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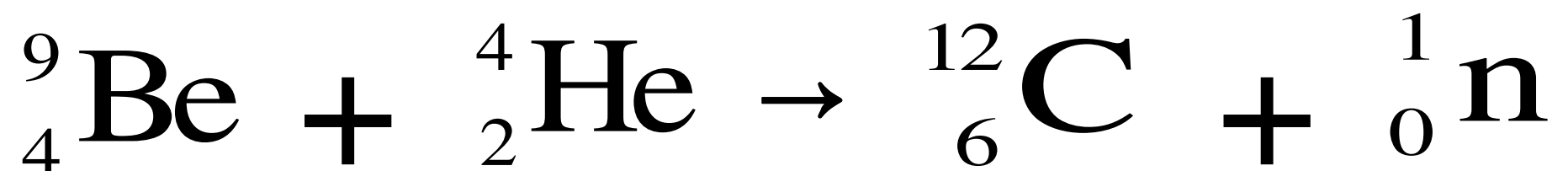
Nuclear Equations



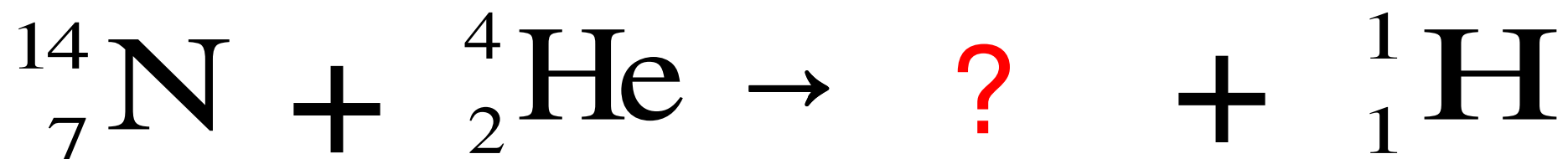
Nuclear Equations



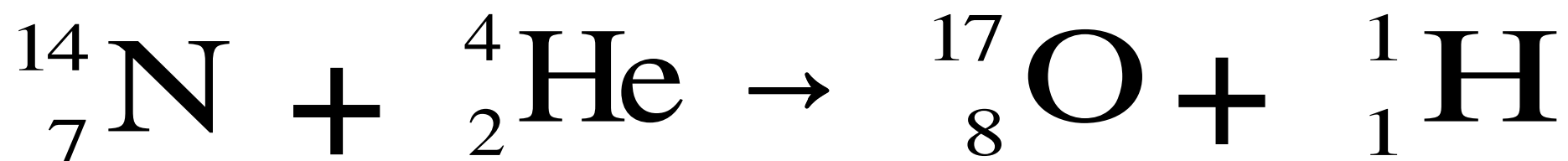
Nuclear Equations



Nuclear Equations



Nuclear Equations



Energy released

$$\Delta E = \Delta m \times c^2$$

Convert:

Atomic Mass Units (amu) into kg

$$1 \text{ amu} = 1.66 \times 10^{-27} \text{ kg}$$

Mass (kg) =

$$\text{mass number} \times 1.66 \times 10^{-27} \text{ kg}$$

(amu)

Energy released

Compare predicted and experimental He masses

Predicted mass =

$$2p^+ + 2n + 2e^- = 4.0320 \text{ amu}$$

Experiment mass = 4.0015 amu

Mass

$$= \Delta m$$

Energy released

$$\Delta E = \Delta m \times c^2$$

Mass → **nuclear binding energy**

$$\begin{aligned} E &= 0.0305 \times 1.66 \times 10^{-27} \times (3 \times 10^8)^2 \\ &= 4.5 \times 10^{-12} \text{ J per atom} \end{aligned}$$

For 1 mole: $E = 2.7 \times 10^9 \text{ kJ}$

Energy released

$$\Delta E = \Delta m \times c^2$$

**Also used to find E released
during any nuclear reaction**

Energy released

$$\Delta E = \Delta m \times c^2$$

Also used to find E released during any nuclear reaction



209.9829

205.9745

4.0026

Need Δm

Energy released

$\Delta m = \text{initial mass} - \text{final mass}$

$= (\text{mass Po}) - (\text{mass Pb} + \text{He})$

$= 209.9829 - (205.9745 + 4.0026)$



209.9829

205.9745

4.0026

$\Delta m = 0.058 \text{ amu}$

Energy released

$$\Delta E = \Delta m \times c^2$$

$$= (0.0058 \times 1.66 \times 10^{-27}) \times (3 \times 10^8)^2$$

$$= 2.89 \times 10^{-13} \text{ J}$$

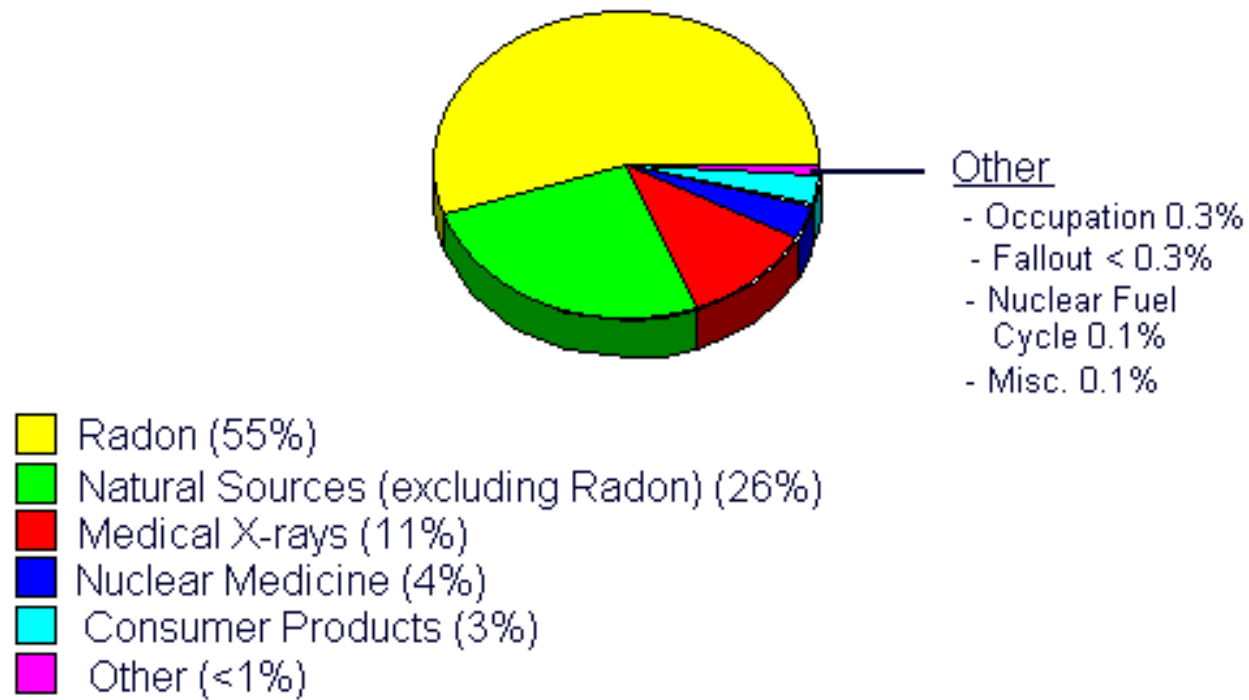


209.9829

205.9745

4.0026

Sources of Radiation Exposure





**Second leading cause of lung cancer in the US.
Causes 1000's of lung cancer deaths per year.**

Radon:
What Is It?



**Colorless, odorless and tasteless
radioactive gas**

In most rocks and soil

Released by decay of uranium

Dispersed outdoors

Possible high levels indoors

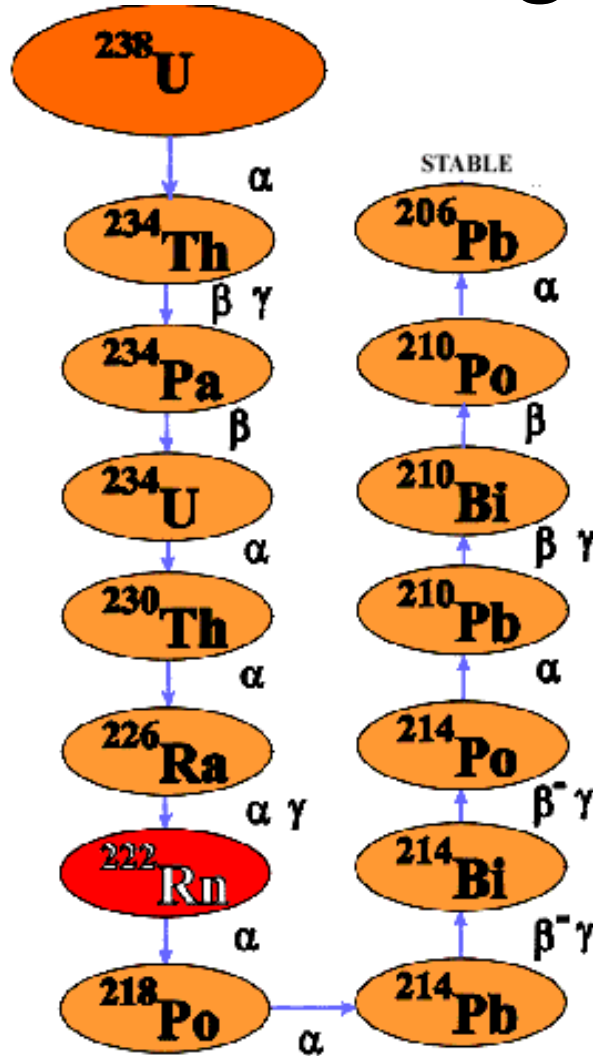


High levels in 15 counties

Madison and Colbert greatest



Where does radon come from ? Why is radon dangerous ?



Radon: gas →

Polonium: solid →

Shielding

Alpha particles

blocked by: 1 cm of air
sheet of paper
human skin

If α -emitting isotopes inhaled or ingested → harmful.

Shielding

Beta particles

requires 1 mm of Al to block

If β -emitting isotopes inhaled or ingested \rightarrow absorbed by bones

Shielding

Gamma and X-rays

need thick concrete or lead to block

most penetrating

severe damage to internal organs

Measuring Radiation

Geiger counter:
radiation ionizes argon \Rightarrow current

Units of radiation

Becquerel (Bq)

SI unit = 1 disintegration/sec

Curie (Ci)

amount of radioactive material
decaying at same rate as 1 g of Ra

= 3.7×10^{10} dis/sec

$1 \text{ Ci} = 3.7 \times 10^{10} \text{ Bq}$

Measure radioactive decay

Units of radiation

Gray (Gy)

1 Gy of radiation deposits 1 J of energy per kg of matter (SI unit)

Rad 1 Rad = 0.01 Gy

Measure equivalent dose

Units of radiation

Sievert (Sv) 1 Sv = 100 rem

Rem

rad equivalent for humans, used to describe biological damage

Measure absorbed dose

Everything is radioactive to some extent

One loaf of bread 70Bq

One adult person 3,000Bq

Time for 50% of an isotope to decay

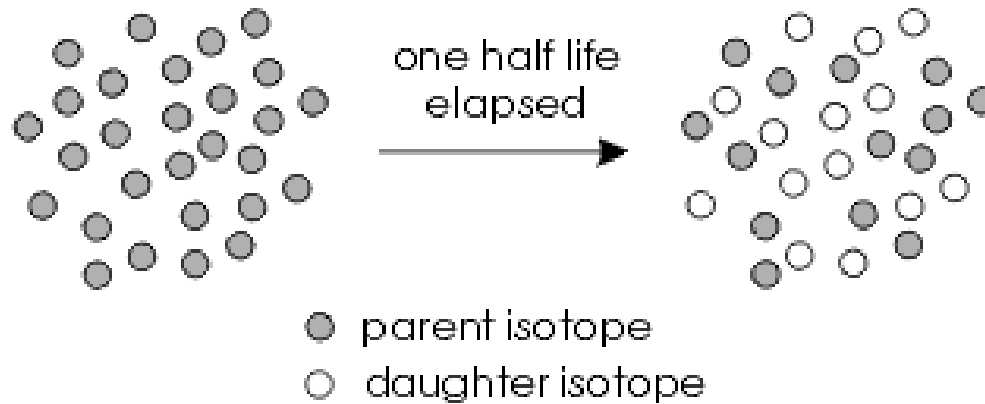
Some Representative Half-Lives

Nuclide	Half-Life ^a	Nuclide	Half-Life ^a	Nuclide	Half-Life ^a
${}^3_1\text{H}$	12.26 y	${}^{40}_{19}\text{K}$	1.25×10^9 y	${}^{214}_{84}\text{Po}$	1.64×10^{-4} s
${}^{14}_6\text{C}$	5730 y	${}^{80}_{35}\text{Br}$	17.6 min	${}^{222}_{86}\text{Rn}$	3.823 d
${}^{13}_8\text{O}$	8.7×10^{-3} s	${}^{90}_{38}\text{Sr}$	27.7 y	${}^{226}_{88}\text{Ra}$	1.60×10^3 y
${}^{28}_{12}\text{Mg}$	21 h	${}^{131}_{53}\text{I}$	8.040 d	${}^{234}_{90}\text{Th}$	24.1 d
${}^{32}_{15}\text{P}$	14.3 d	${}^{137}_{55}\text{Cs}$	30.23 y	${}^{238}_{92}\text{U}$	4.51×10^9 y
${}^{35}_{16}\text{S}$	88 d				

^as, second; min, minute; h, hour; d, day; y, year.

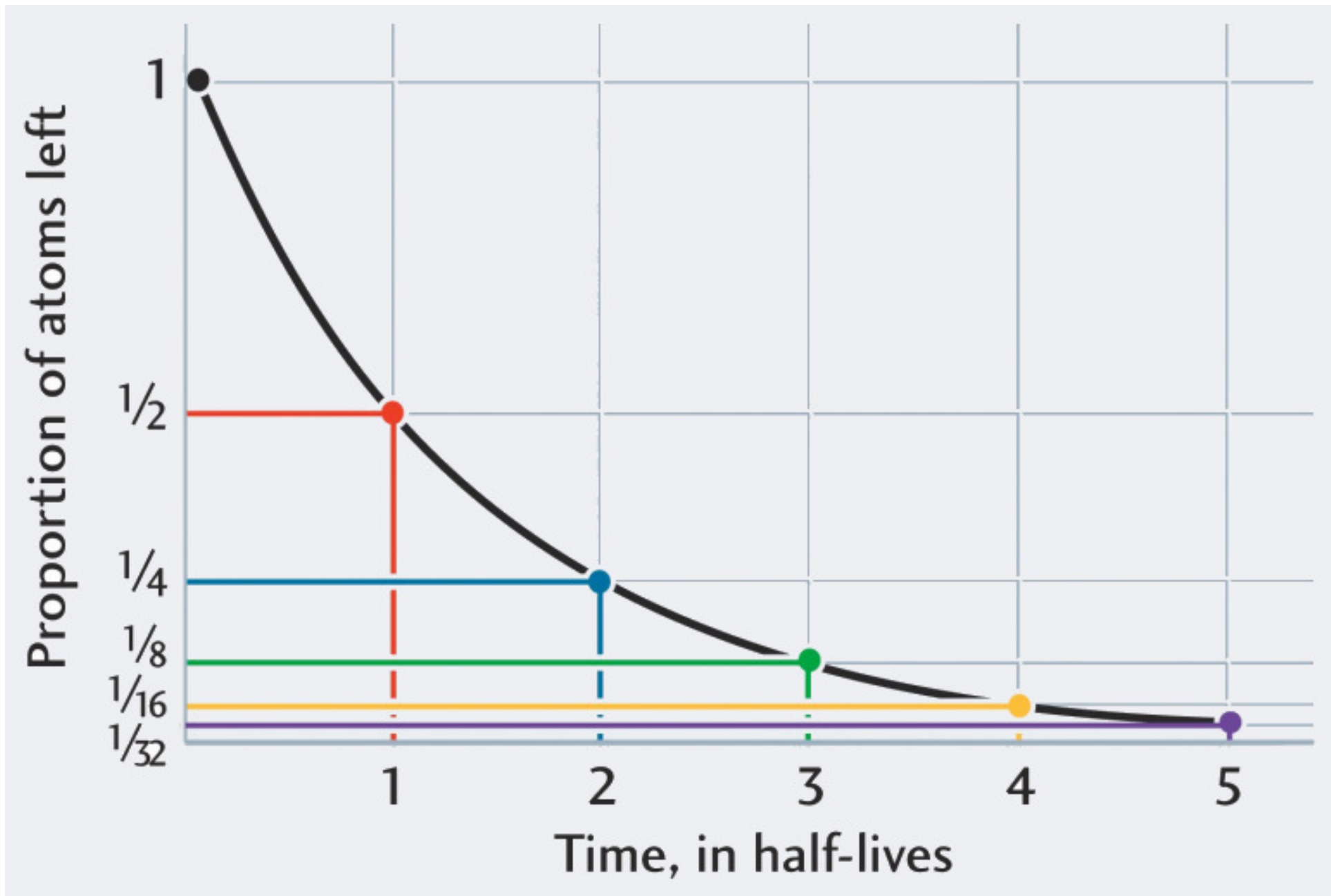
Radioactive Decay

Half Life



Decay of ^{14}C





Half-life

First-order reaction

$$t_{1/2} = \frac{0.693}{k}$$

$$\log \frac{C_0}{C_t} = \frac{kt}{2.303}$$

k = rate constant

C₀, C_t = initial, final amounts

$$\log \frac{C_0}{C_t} = \frac{kt}{2.303}$$

- how much isotope left after time, t
-
- time taken for % of isotope to decay

Example:

Half-life of ^{63}Ni is 100 years.

If you had 100 g of ^{63}Ni , how much would be left after 250 years?

$$\log \frac{C_0}{C_t} = \frac{kt}{2.303}$$

$$\log \frac{C_0}{C_t} = \frac{kt}{2.303}$$

$$k = ?$$

$$C_0 = 100 \text{ g}$$

$$C_t = ?$$

$$t = 250 \text{ yr}$$

$$\log \frac{C_0}{C_t} = \frac{kt}{2.303}$$

$$k = ?$$

$$C_0 = 100 \text{ g}$$

$$C_t = ?$$

$$t = 250 \text{ yr}$$

$$t_{1/2} = \frac{0.693}{k}$$

$$k = 0.0693 \text{ yr}^{-1}$$

$$C_t = 17.7 \text{ g}$$

Radiation and Health

Factors influencing degree of exposure

Half-life shorter half-life materials decay faster \Rightarrow greater damage

Type of radiation some worse than others

Distance from source Intensity $\propto 1/\text{distance}^2$

Time of exposure cumulative

Uses of radiation

- **Age determination (dating) of minerals/fossils**
- **Cancer treatment**
- **Tracers & Imaging**
- **Nuclear power**
- **Food irradiation**

Radioactive Dating

two general types

geochronology

long half-life isotopes in minerals

carbon dating

radioactive C-14 in formerly living objects

Radiocarbon Dating

Two carbon isotopes found in nature:

carbon-12 **non-radioactive**

carbon-14 **radioactive**

Radiocarbon Dating

**In living tissue ratio
C-14 to C-12 is 1:10¹²**

**After death, no more C-14 taken in;
ratio changes**

Decreases by 50% every 5,730 yrs

Good for fossils, wood, textiles

Nuclear Power

Energy obtained 2 ways

Fission: splitting large atoms
(power plants; bombs)

Fusion: joining small atoms (sun)

**Nuclear energy provides about
20 % of US electricity**

**103 nuclear reactors with
operating licences in 31 states**

25 % of electricity in Alabama

Isotopes of Uranium



Pitchlend

Nuclear Power

Uranium-235 used as reactor fuel
Produces nuclear chain reaction by fission



Nuclear Power

Fission generates heat

Heat boils water

steam



turbines



electricity

**Two sections U-235
forced together**

**Creates
supercritical mass**

When just enough fissions occur to keep the chain reaction going → **critical reaction** → nuclear power

When excess neutrons produced fission rate keeps increasing → **supercritical reaction** → nuclear explosion

Little Boy and Fat Man

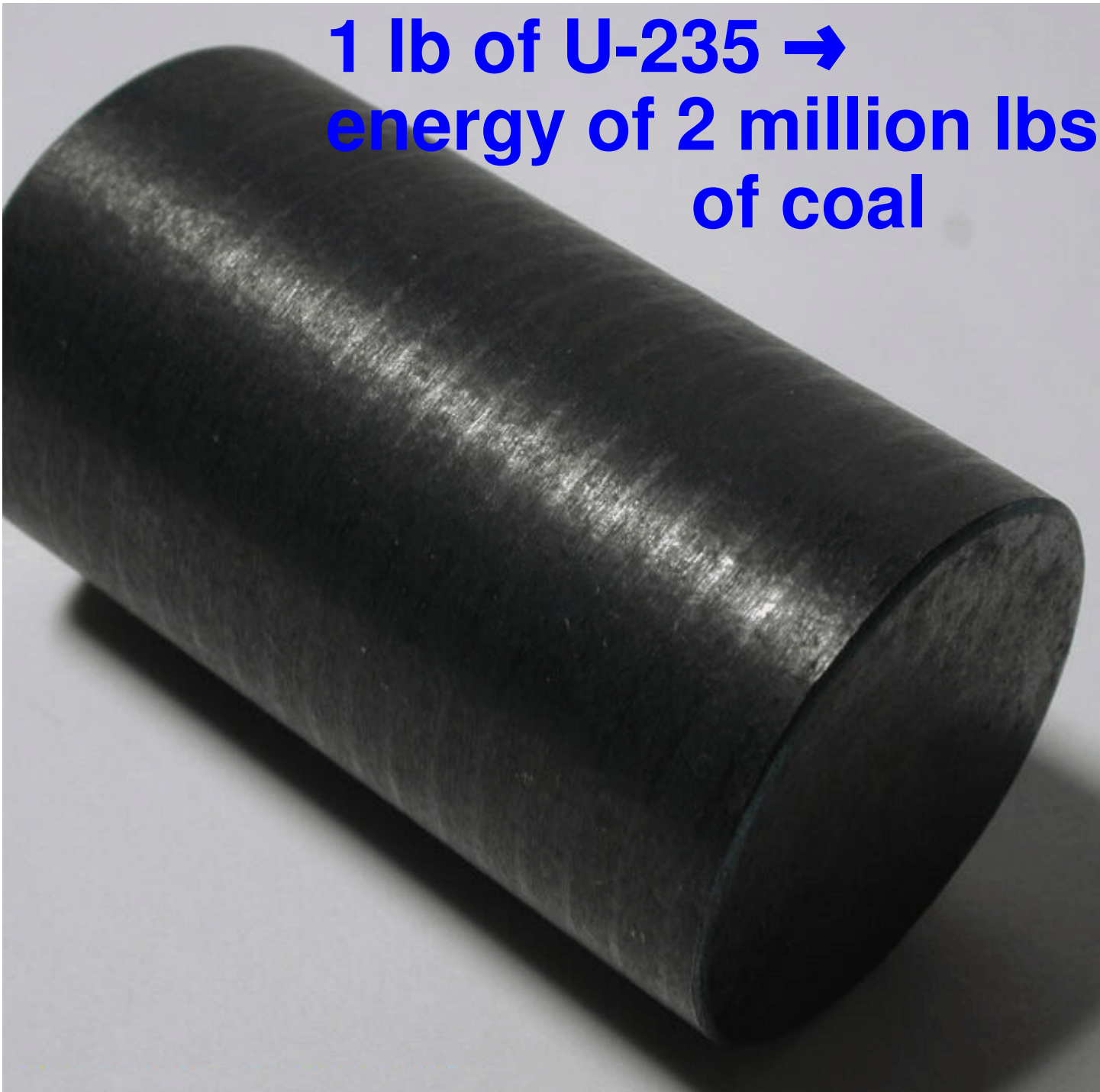
August 6th, 1945: 8:15am.

**Enola Gay releases "Little Boy" at altitude 31,500'.
Preset to explode at altitude 2000' above Hiroshima.
75,000 people killed immediately.
48,000 buildings destroyed.**

August 9th, 1945: 10:00am.

**Bock's Car releases "Fat Man" over Nagasaki.
35,000 people killed immediately.**

**1 lb of U-235 →
energy of 2 million lbs
of coal**



Uranium tipped tank penetrating munitions

Why use uranium?

- 1. Very, very dense**
- 2. Very hard**
- 3. Pyrophoric**

Health Effects

Two classes

Somatic (whole body) damage

- ▶ Chronic (cancer)
- ▶ Acute (acute radiation syndrome)

Genetic damage

- ▶ Alters genetic material
- ▶ Mutations in offspring

Cadmium control rods absorb excess neutrons to keep reactor below supercritical

Fusion

**Solar output =
E of burning 1500 lb coal/hour
for each sq² of solar surface**

Fusion

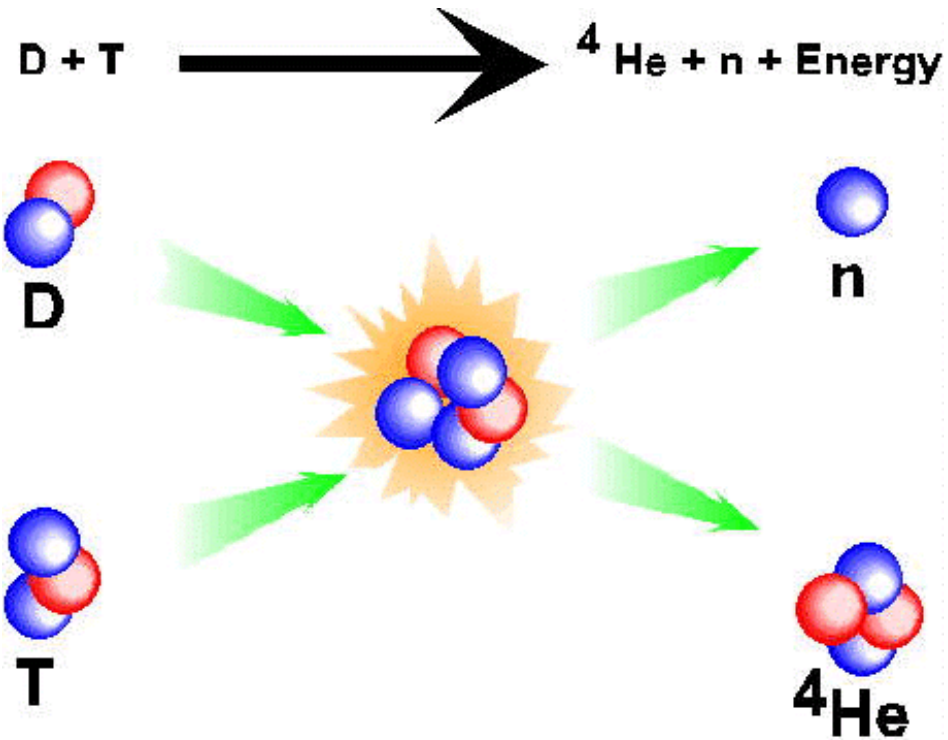
**1919: Sir Arthur Eddington
showed $H \rightarrow He$
provided Sun's energy**

Fusion

4 million metric tons H per sec

100 billion H-bombs each sec

Fusion



Deuterium & Tritium come from water

